Pilot Data Analysis of PDS Geosciences Data in the Cloud with the Java CRISM Analysis Tool (JCAT)



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WHY ANALYZE DATA ON THE CLOUD?

Doing heavy computing on the Cloud reduces requirements for the user's own computer, which lowers hardware costs. Short-term use of a cloud VM to do an analysis is often relatively cheap (see "Methodology").

Local computer requirements change for these common processing bottlenecks:

- Storage space: doesn't need to be enough to fit data set under analysis
- CPU: doesn't need to be as powerful
- RAM: doesn't need to be large enough to run the analysis software

Other benefits:

- Faster to acquire data, if data supplier (e.g., PDS Geosciences Node) stores data in the same cloud data center (see Figure 1 in below section)
- Provides temporary processing and storage resources
- · Scalable cloud processing to meeting application requirements
- Data supplier also benefits from avoiding data egress fees (i.e., a charge incurred upon download), which are often a significant portion of costs of data storage on the cloud

LOCAL VS. CLOUD DATA ANALYSIS



Figure 1: Some differences between local and cloud data analysis of PDS Geosciences Node (GEO) data. In local analysis, large files must be transferred to the user's local machine across the public internet, potentially leading to large delays. In cloud analysis, however, large files are transferred within a cloud data center's local network and so can be ready for analysis sooner. Only the smaller result files need to be transferred back to the user's local machine.

METHODOLOGY

In this pilot project, we analyzed data from the Compact Reconnaissance Imaging Spectrometer for Mars (CRISM) hyperspectral instrument using the Java CRISM Analysis Tool (JCAT) (http://crism.jhuapl.edu/JCAT/) available from JHU/APL.

The techniques and tools used in this project, as well as the rationale for using the cloud, are applicable to many other PDS datasets and the tools for analyzing them.

We used a Windows Server 2019 VM in the Microsft Azure Central US zone with configuration Standard_D2s_v3:

- 2 virtual CPUs
- 8 GB of RAM
- ~100 GB of virtual hard-drive space available
- Windows license

Our costs: \$0.2110/hour. Expected time to analyze 10 CRISM scenes with JCAT as described in "Example CRISM Analysis" is 7 hours, making the total analysis cost \$1.48. (Note: this cost may vary between users, depending on relationships with the cloud provider.)

Steps before the analysis:

- From Azure cloud management web portal, selected desired VM configuration and logged into it with remote desktop
- Transferred installers for non-OS software we needed (Python, Java Runtime Environment, JCAT itself) to the VM using our remote desktop client's "local resources" feature
- Installed the above tools on the VM
- Found CRISM products for analysis using Geosciences Node's Mars Orbital Data Explorer (https://ode.rsl.wustl.edu/mars/)
- Copied download links for these files to our VM & and input them to a Python script that used them to locate & download the files from the archive cloud copy

EXAMPLE CRISM ANALYSIS



Figure 2: JCAT's initial view of CRISM scene HRL000040FF's data upon loading it. The displayed IOF spectrum was chosen from the candidate carbonate-bearing terrain. Vertical lines show the spectral bands used to display the RGB false color images in the upper left. Carbon dioxide band triplet at ~2 micrometers is evident, along with the deep 3 micrometer band and thermal rise at longer wavelengths. North is toward the bottom of the image views.

To illustrate this proof-of-concept, we analyzed CRISM scene HRL000040FF, which covers the western side of Jezero Crater, including the Mars 2020 rover mission (Perseverance) landing error ellipse, the delta, and the carbonate-bearing terrains up against the western Jezero Crater wall [1]. Our analysis:

- Loaded IOF data for the scene into JCAT (see Figure 2 above)
- Divided the spectrum by a volcano scan observation to remove the gas bands
- Identified diagnostic 2.3 and 2.5 micrometer absorptions in an area of the scene, which demonstrates the presense of carbonate-bearing terrain (see FIgure 3 below)
- · Exported these results from JCAT as PNG images
- · Transferred our results back to our local machine via remote desktop client



Figure 3: IOF spectrum shown after dividing by a volcano scan to remove gas bands, with the X and Y ranges of the spectral plot configured to better show absorption features at 2.3 and 2.5 micrometers associated with one or more carbonate bearing minerals.

References: [1] Goudge T. A. et al. (2015) JGR Planets, 120, 775-808.

FUTURE PLANS

Our possible future directions involve better support for data users who want to use this workflow for their own planetary data analysis, by making the process simpler. Possibilities include:

- "Download from Cloud" option in PDS Geosciences search tools, to be used directly from user's VM (so custom download script is not necessary)
- Cloud data mounted in VM as read-only disk
- Pre-made VM image with pre-installed basic enabling software (Python, Java Runtime Environment, etc.)
- Online documentation, including examples
- Support for similar procedures with cloud VMs running Linux instead of Windows
- Tools from proprietary software vendors that allow users to run licenced software, often without licences for certain subsets of use-cases

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[1] Goudge T. A. et al. (2015) JGR Planets, 120, 775-808.